

INAF



ISTITUTO NAZIONALE DI ASTROFISICA
OSSERVATORIO ASTROFISICO DI TORINO

Cold Jupiters and short-period small planets: Friends, foes, or indifferent?

A.S. Bonomo, L. Naponiello, E. Pezzetta, A. Sozzetti,
D. Gandolfi, R. Wittenmyer, and M. Pinamonti

Bonomo et al. (2025), A&A, 700, A126

51 Peg b - Cool giant planets and their systems
Observatoire de Haute Provence (6-10 October, 2025)

Cold Jupiters (CJs) and small planets (SPs) are friends?

f_{CJ} : frequency of CJs around solar-type stars, regardless of the presence/absence of SPs

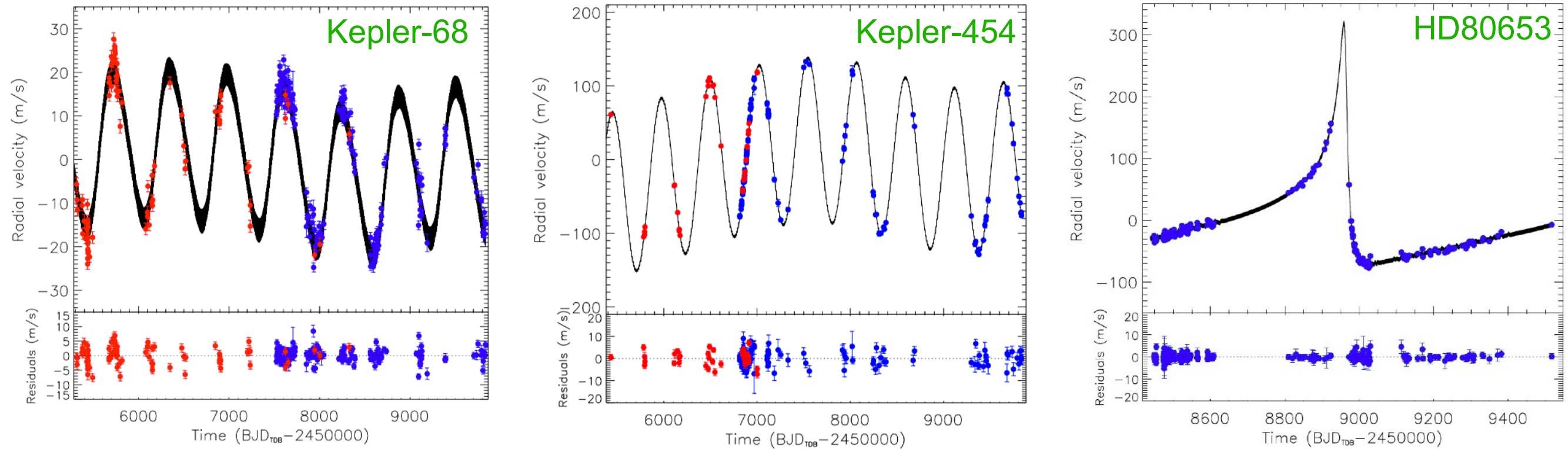
$f_{\text{CJ}|\text{SP}}$: frequency of CJs around solar-type stars with SPs (inside-out approach)

Yes: $f_{\text{CJ}|\text{SP}} = 38 \pm 7 \%$ $\Rightarrow$ excess of CJs in SP systems (Bryan+2019; Zhu & Wu 2018)

Maybe: $f_{\text{CJ}|\text{SP}} = 29 (-11 + 14) \%$ $\Rightarrow$ possible excess of CJs in SP systems (Rosenthal+2022; Van Zandt+2025)

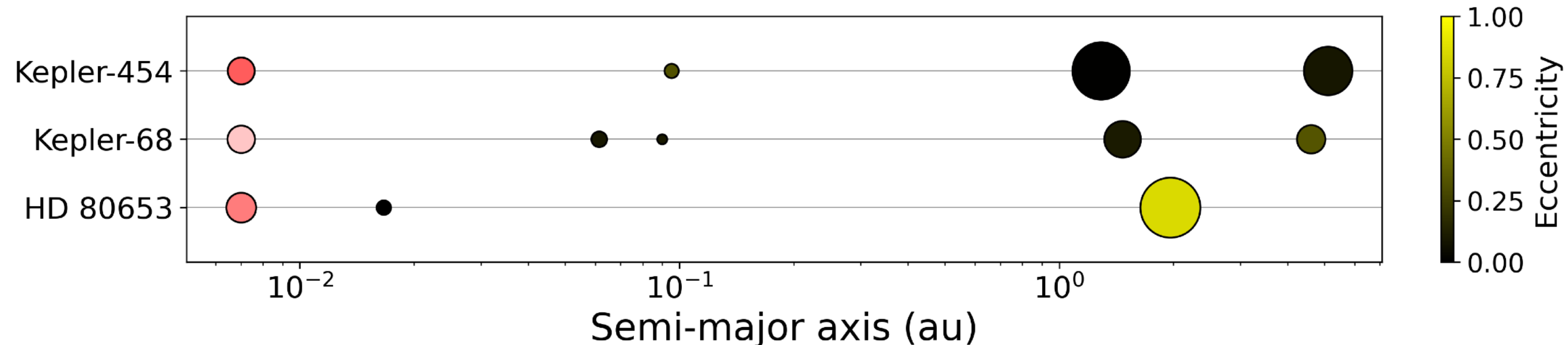
No: $f_{\text{CJ}|\text{SP}} = 9 (+7 -3) \%$ no excess of CJs in SP systems (Bonomo+2023) from the HARPS-N survey of 38 Kepler and K2 SP systems: 3 instead of 12 ± 2 SP systems with CJs

SP systems with CJs found by the HARPS-N Collaboration

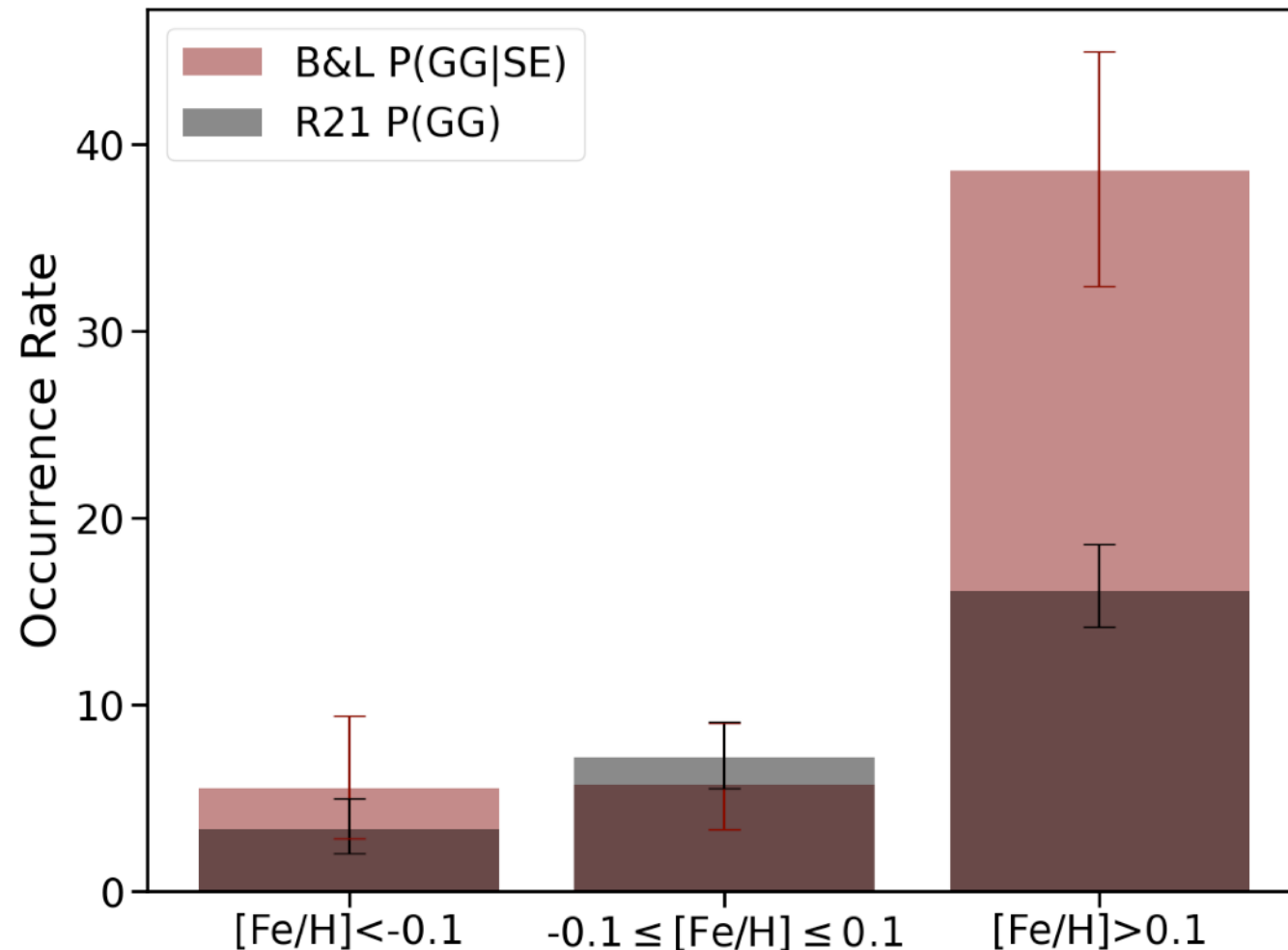


Blue circles: HARPS-N@TNG data
Red circles: HIRES@Keck data (Weiss+2024)

(Bonomo, Dumusque, Massa et al. 2023)



CJs and SPs are friends only at $[\text{Fe}/\text{H}] > 0.1$?



(Bryan & Lee 2024)

By averaging over all metallicities

$$f_{\text{CJ}|\text{SP}} = 15 \text{ (+3 -2) \%}$$

Bryan & Lee 2024 ($\langle \text{Fe}/\text{H} \rangle = -0.001$ dex)

VS

$$f_{\text{CJ}|\text{SP}} = 9 \text{ (+7 -3) \%}$$

Bonomo+2023 ($\langle \text{Fe}/\text{H} \rangle = -0.065$ dex)

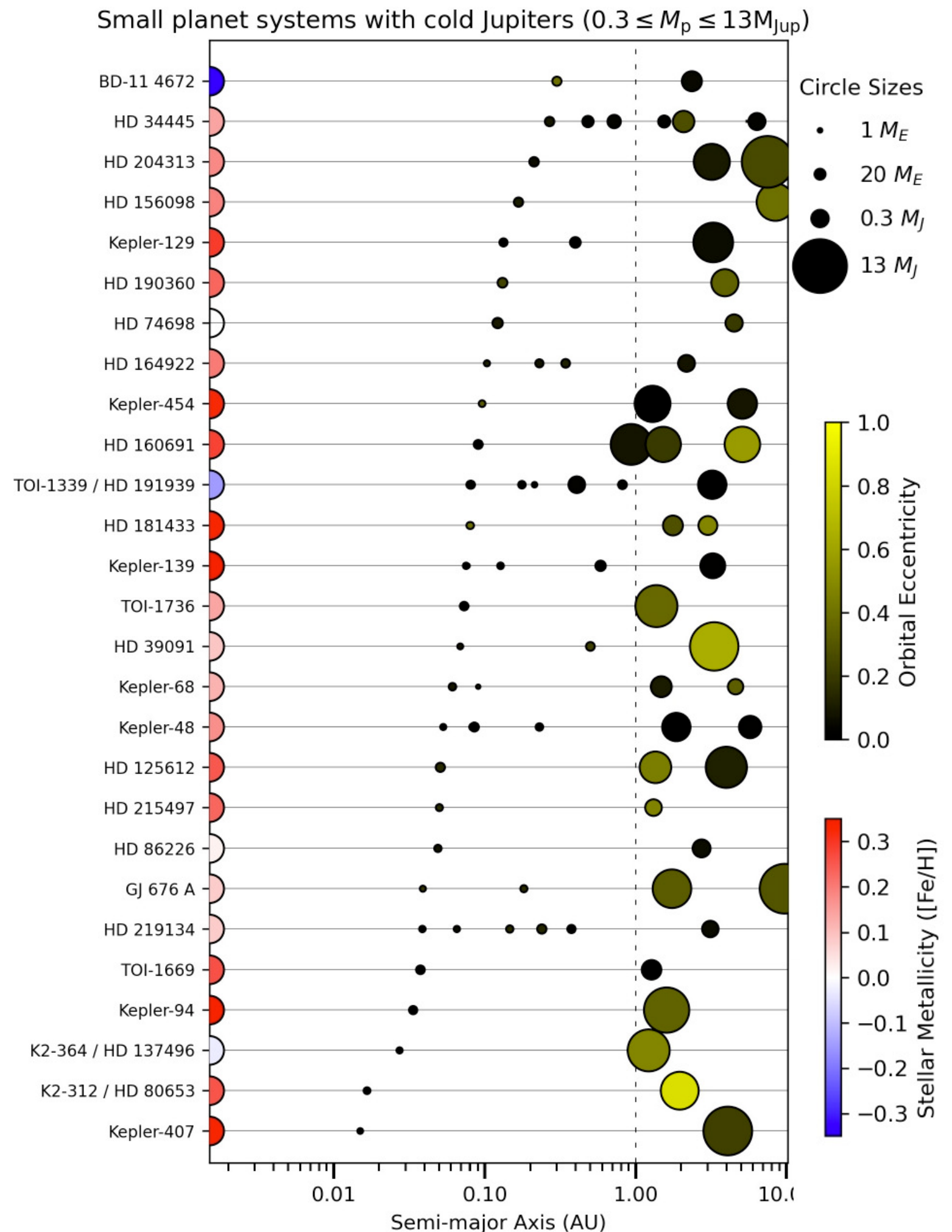
- SP sample (including “mixed systems”): **184 SP systems** (109 transit + 75 RV) with at least one SP with $1 < R_p < 4 R_\oplus$; $1 < M_p < 20 M_\oplus$, $n_{\text{RV}} > 20$, $\Delta T > 1$ yr, $M_\star > 0.6 M_\odot$
- **Comparison sample of 562 stars** from the CLS (California Legacy Survey, Rosenthal et al. 2021) RV survey
- Homogeneous completeness (detection sensitivity) from RV formal uncertainties (no jitter)

Our samples

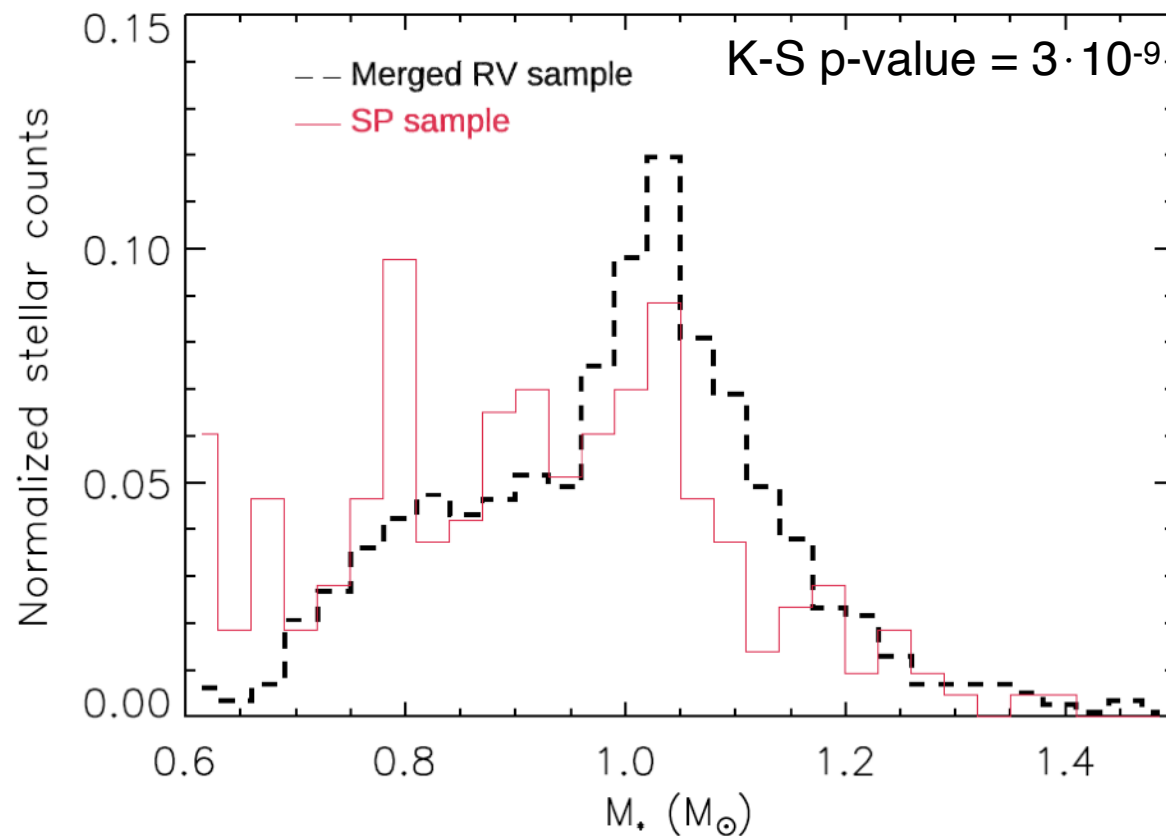
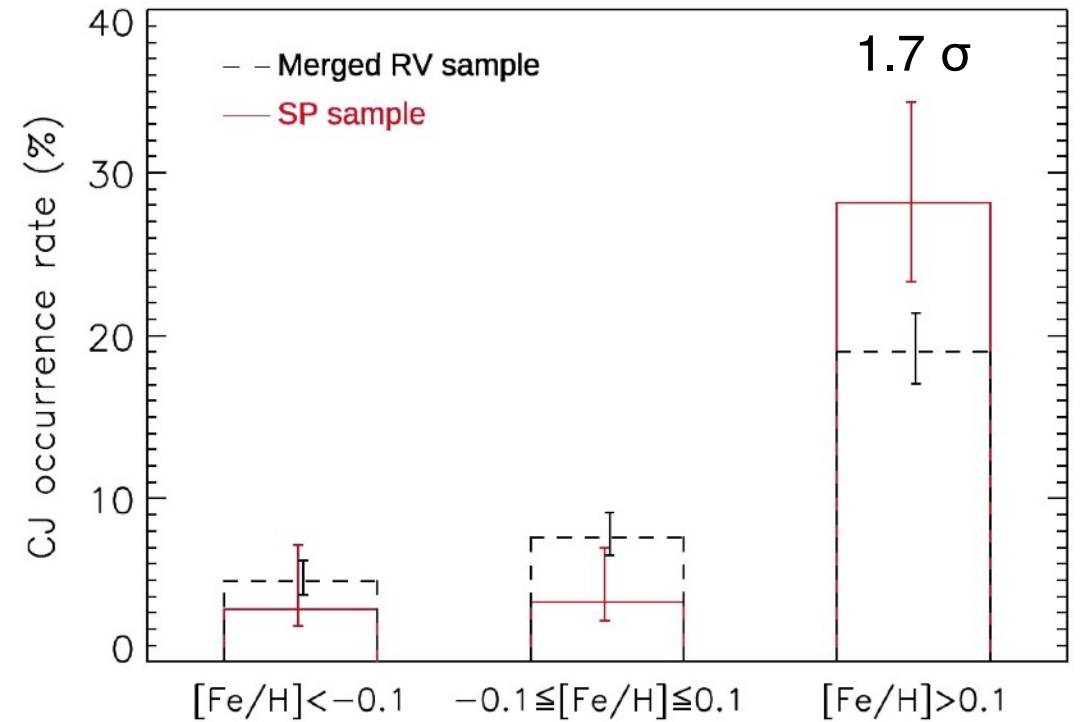
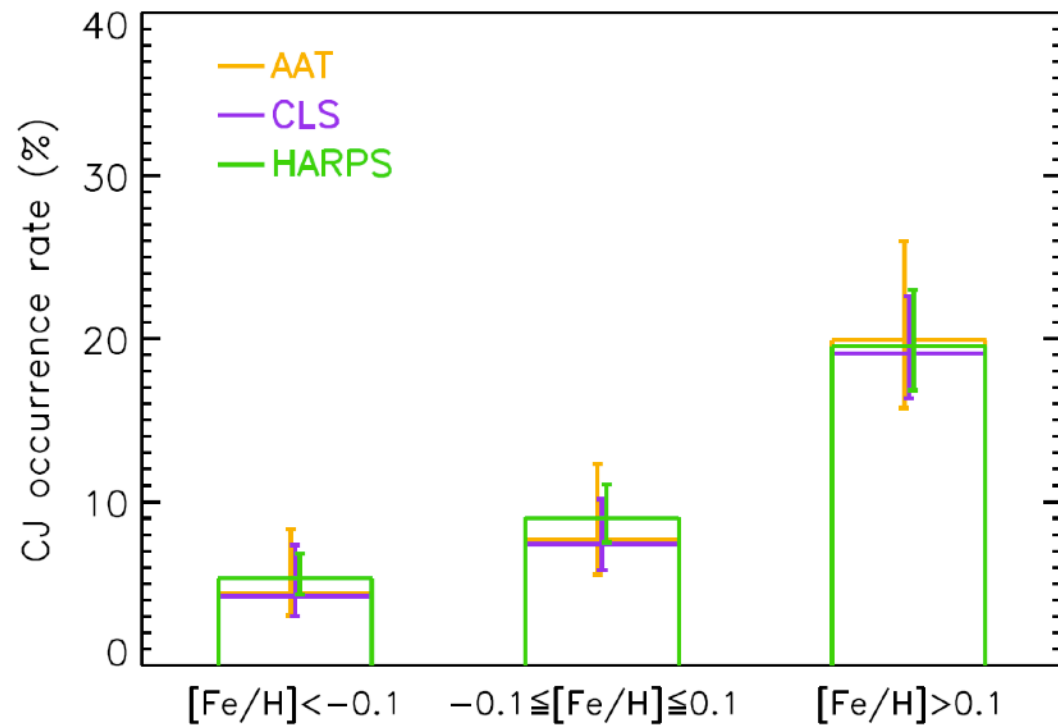
- SP sample (with and without “mixed systems”): **217 SP systems** (134 transit + 83 RV) with at least one SP with $1 < M_p < 20 M_\oplus$, $n_{RV} > 20$, $\Delta T > 1$ yr, $M_\star > 0.6 M_\odot$

- **Comparison sample of 1167 stars** from the HARPS, CLS, and AAT surveys

- Homogeneous completeness (detection sensitivity) from RV formal uncertainties (no jitter)

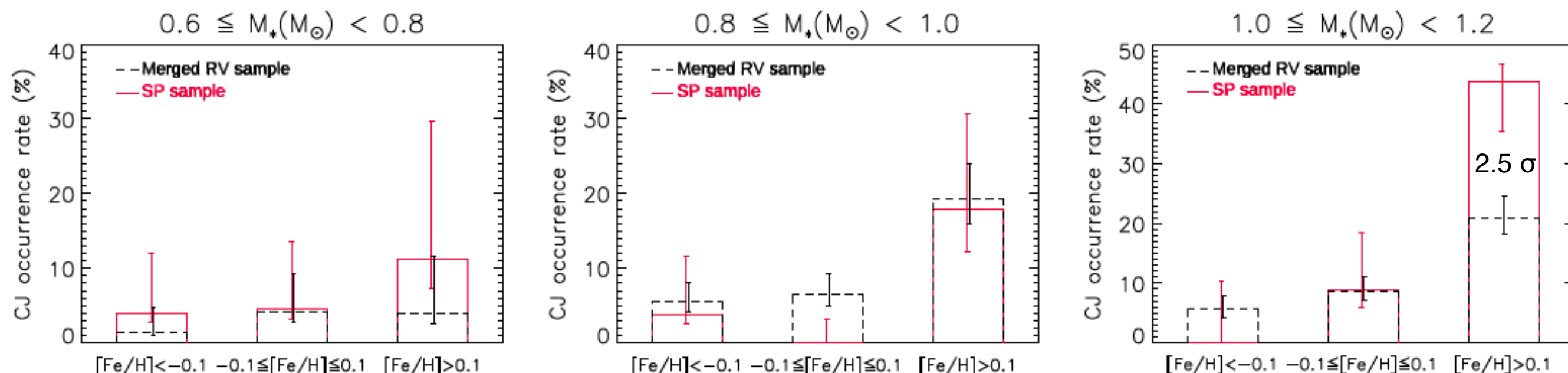


CJs and SPs are friends only at $[\text{Fe}/\text{H}] > 0.1$?



Bias on the $M_\star$ distribution of the SP sample due to the RV and/or transit selection effects.

CJs and SPs are friends only at $[\text{Fe}/\text{H}] > 0.1$ and $M \geq 1 M_{\odot}$?



$$f_{\text{CJIS}} = 11.1 \text{ } (+2.5 \text{ } -1.8) \% \quad \text{vs} \quad f_{\text{CJ}} = 9.8 \text{ } (+0.9 \text{ } -0.8) \%$$

Consistent results on the dependence of f_{CJIS} on $[\text{Fe}/\text{H}]$ and $M_{\star}$ were independently found by Bryan & Lee 2025.

N.B.: f_{CJIS} and f_{CJ} may slightly change if long-term RV trends turn out to be CJs and/or $M_p \sin i$ of CJs from Gaia astrometry turn out to be $M_p > 13$ or $20 M_{\text{Jup}}$. Moreover, they must be updated as more and more well-characterized systems are available.

CJs and SPs are friends or foes?

- **Anticorrelation**: CJs as dynamical barriers to sub-Neptune inward migration (e.g., Izidoro+2015) or hindrance to SP formation inside the water snowline (Lambrechts+2019)
- **Weak or no correlation** from planetesimal (Schleckler+2021) or pebble (Danti+2025) accretion
- **Strong correlation** for low efficient gas contraction rate (Bitsch & Izidoro 2023) or planetesimal snowplow-like process due to CJ-induced secular resonances (Best+2024, 2025).

CJs and SPs are friends or foes?

- **Anticorrelation**: CJs as dynamical barriers to sub-Neptune inward migration (e.g., Izidoro+2015) or hindrance to core formation inside the water snowline (Lambrechts+2019) 
- **Weak or no correlation** from planetesimal (Schleckler+2021) or pebble (Danti+2025) accretion
- **Strong correlation** for low efficiency of contraction rate (Bitsch & Izidoro 2023) or planetesimal snowplow-like process due to CJ-induced secular resonances (Best+2024, 2025). 

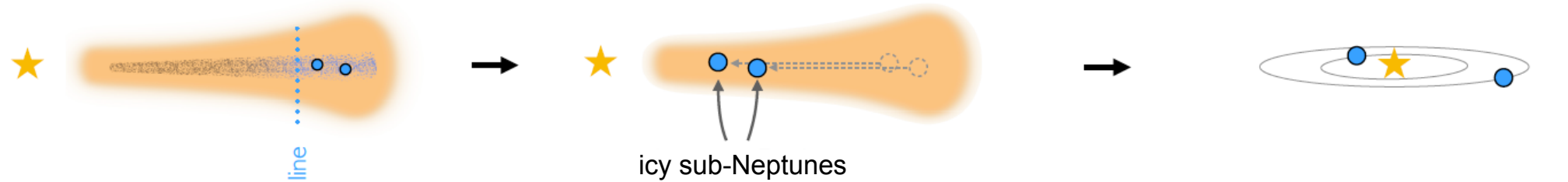
CJs and SPs are friends or foes?

- **Anticorrelation**: CJs as dynamical barriers to sub-Neptune inward migration (e.g., Izidoro+2015) or hindrance to core formation inside the water snowline (Lambrechts+2019) 
- **Weak or no correlation** from planetesimal (Schleckler+2021) or pebble (Danti+2025) accretion
- **Strong correlation** for low efficient gas contraction rate (Bitsch & Izidoro 2023) or planetesimal snowplow-like process due to CJ-induced secular resonances (Best+2024, 2025). 
 - at $[\text{Fe}/\text{H}] > 0.1$ and $M \geq 1 M_{\odot}$

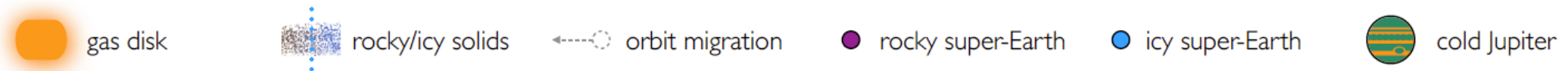
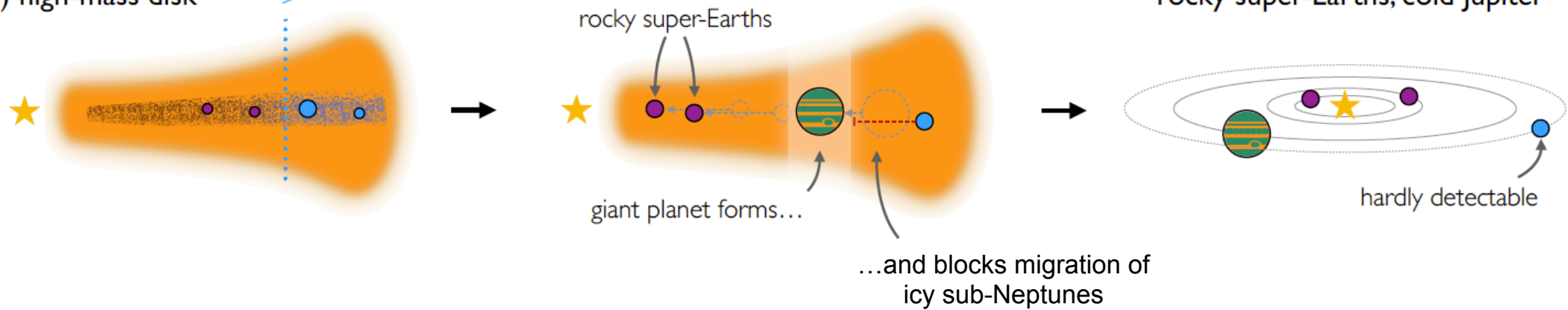
Any strong architecture-composition link?

Generation 3 Bern Planet Population Synthesis (Schlecker et al. 2021)

a) intermediate-mass disk

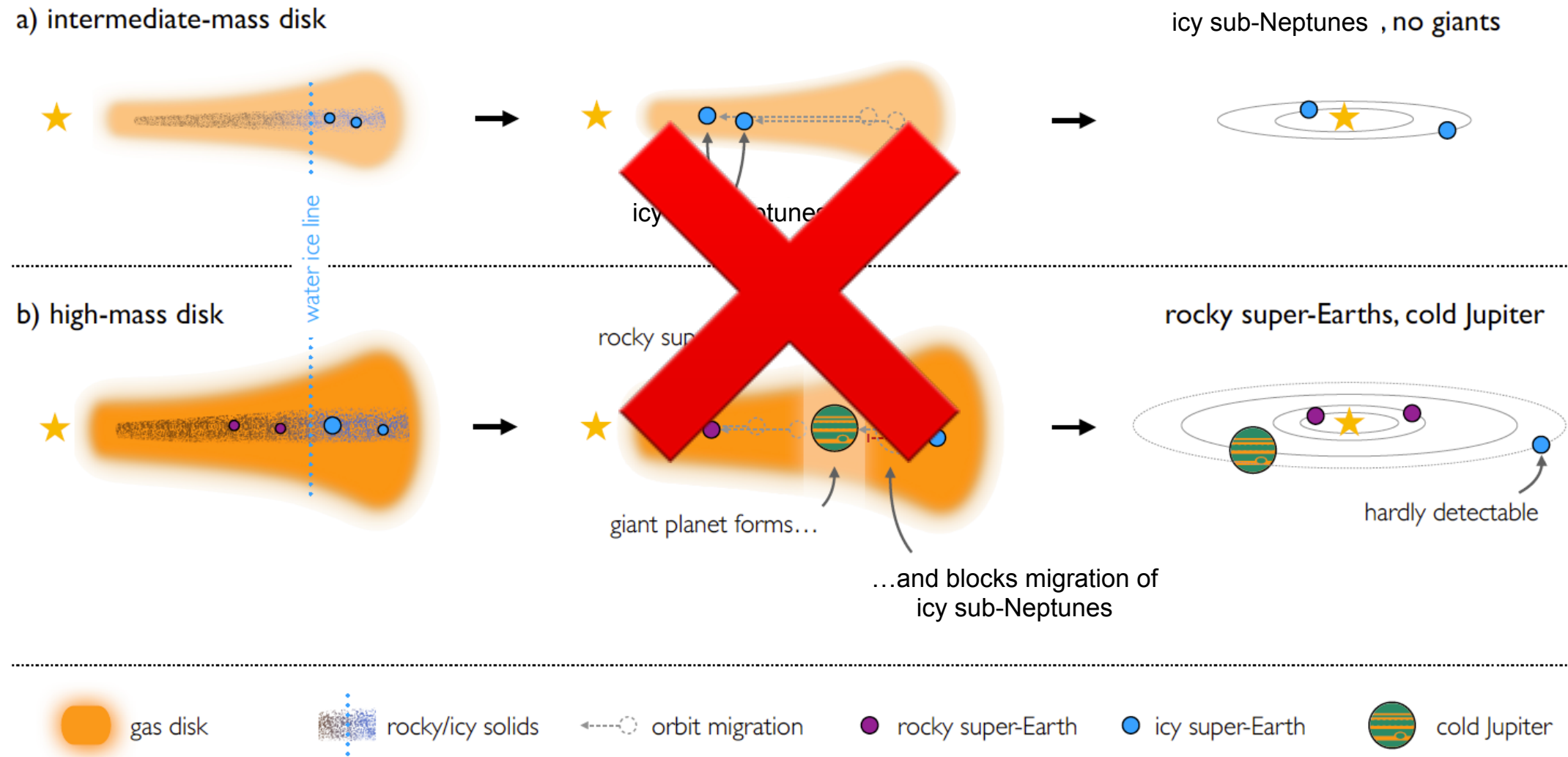


b) high-mass disk



Any strong architecture-composition link?

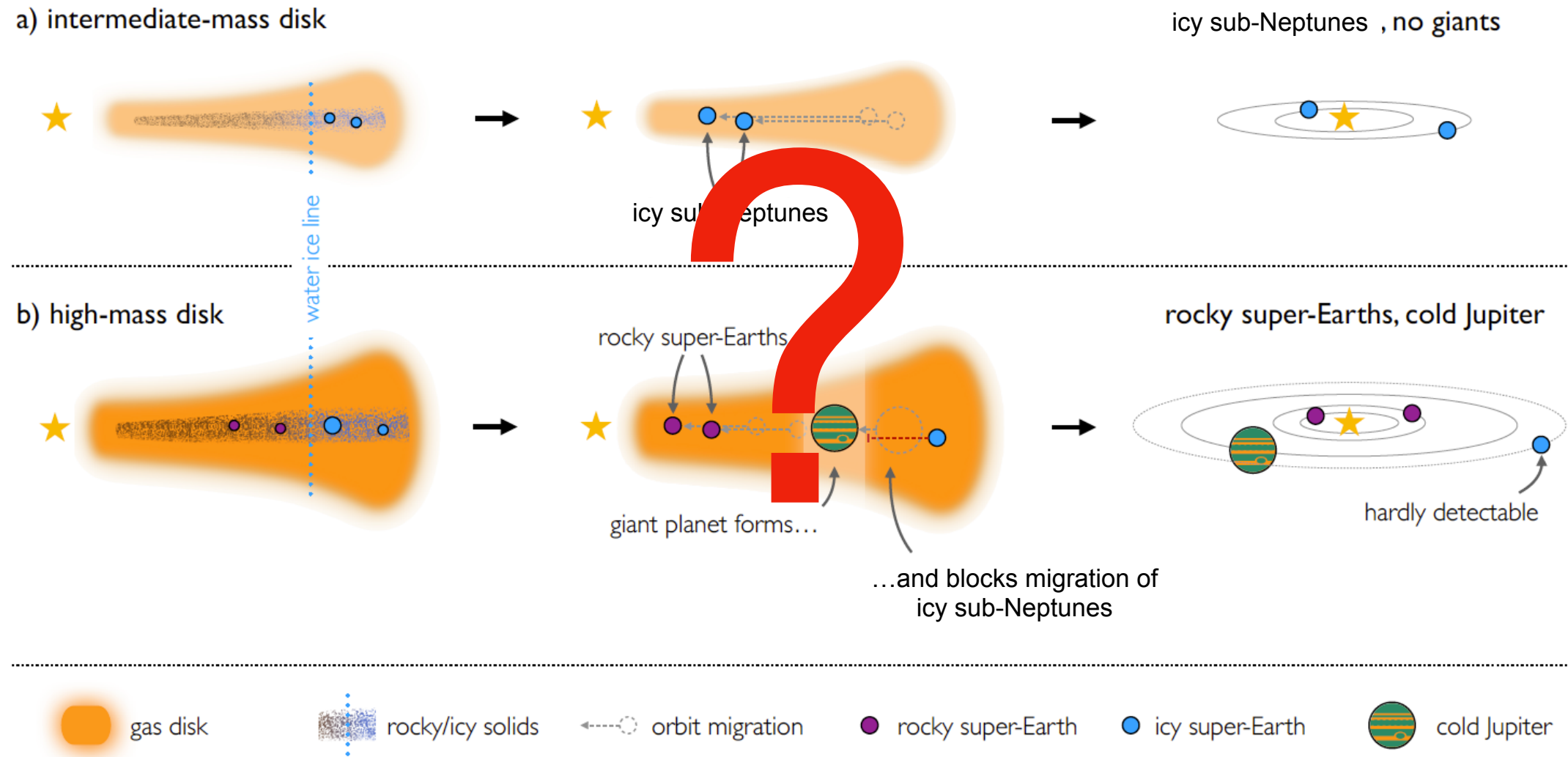
Generation 3 Bern Planet Population Synthesis (Schlecker et al. 2021)



$$f_{\text{rocky,w/CJ}} = 33 \text{ } (+13 \text{ } -9) \%$$

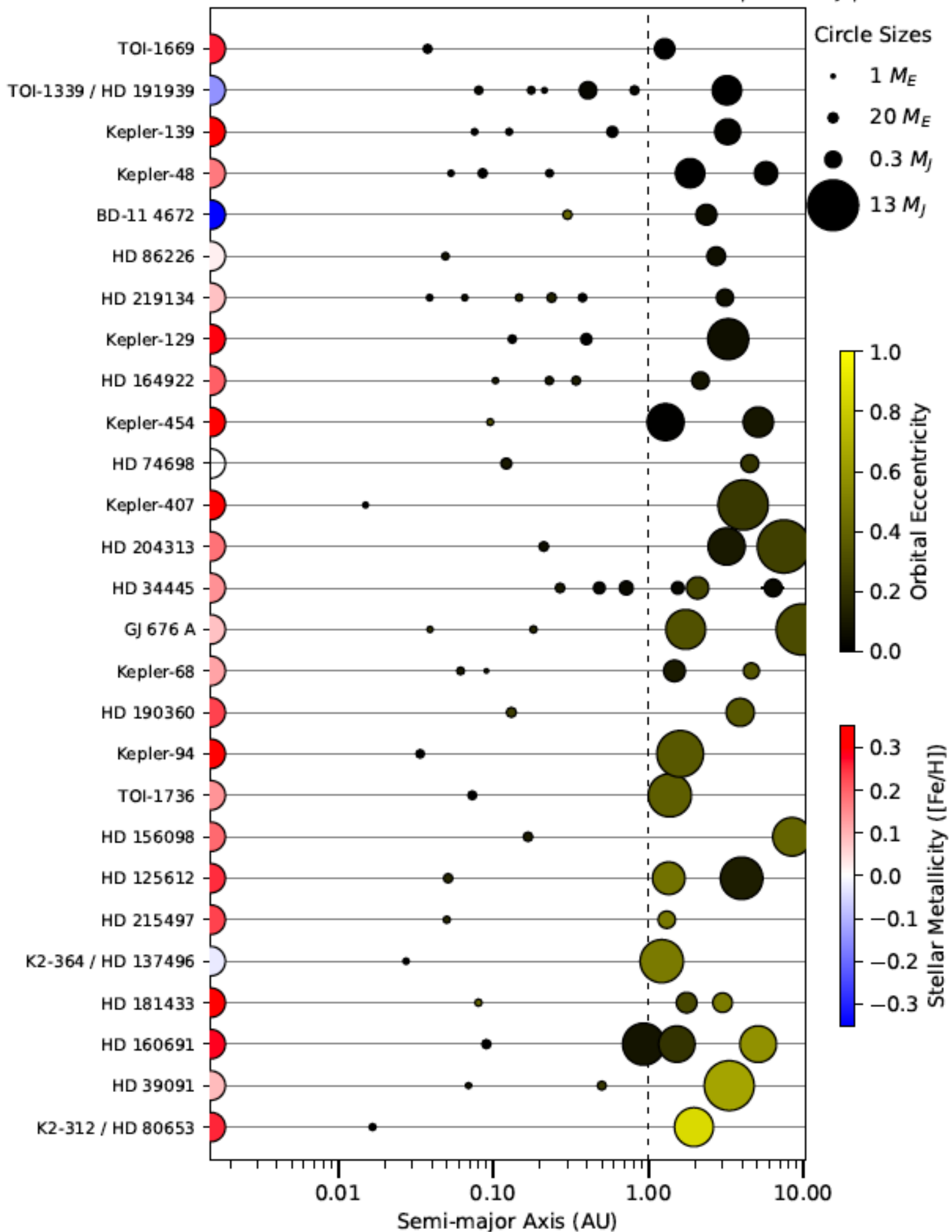
Any strong architecture-composition link?

Generation 3 Bern Planet Population Synthesis (Schlecker et al. 2021)



$$f_{\text{rocky,w/CJ}} = 33 \text{ } (+13 \text{ } -9) \%$$

Small planet systems with cold Jupiters ($0.3 \leq M_p \leq 13M_{\text{Jup}}$)



SP systems are NOT indifferent to CJs

The architecture of SP systems can be greatly influenced by CJs (e.g., dynamical instabilities in inner regions and/or SP inclination excitation)

On the shoulders of giants

Cold Jupiters and their role in the formation of inner low-mass planets. From theory to observations

25-26-27 March 2026

Turin



Information on registration and abstract submission will be soon available

Invited speakers: *Eleonora Alei, Bertram Bitsch, Timothy Brandt, Marta Bryan, Juan Cabrera, Gael Chauvin, Joao Faria, Jonathan Horner, Anne-Marie Lagrange, Gijs Mulders, Luca Naponiello, Matteo Pinamonti, Sean Raymond, Judah Van Zandt, Lauren Weiss.*